

gravitation, problem set #2

1) We know the radius of the earth, $R_E = 6.4 \times 10^3$ km, and the gravitational field on the surface, $g = 9.8 \text{ m/s}^2$. Also we know moon goes around the earth once in $T = 29$ days and its orbit is almost a circle.

- Find the mass of the earth.
- How far is the moon from the earth?
- We see the moon on the sky with $\theta = 0.5^\circ$ angular diameter. Use previous part to find moon's radius.
- Find moon's velocity relative to the earth.
- With more accurate observation, we see that earth is rotating around the center of mass of the earth-moon system which is located at $R_{CM} = 4.7 \times 10^3$ km from center of the earth. Find mass of the moon.

$$a) \quad g = \frac{GM_E}{R_E^2} = 9.8 \text{ m/s}^2 \Rightarrow M_E = \frac{gR_E^2}{G}$$

$$b) \quad r^3 = GM \frac{T^2}{4\pi^2} \rightarrow r = \left[\frac{GMT^2}{4\pi^2} \right]^{1/3}$$

$$c) \quad \tan \frac{\theta}{2} = \frac{R_{Moon}}{r}, \quad \theta \ll 1 \rightarrow \theta = \frac{2R_{Moon}}{r} \rightarrow R_{Moon} = \frac{r\theta}{2}$$

$$d) \quad v = \frac{2\pi r}{T}$$

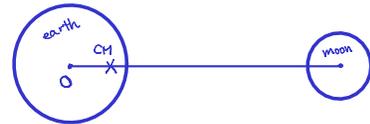
or

$$\frac{v^2}{r} = \frac{GM_E}{r^2} \Rightarrow v = \sqrt{\frac{GM_E}{r}}$$

$$e) \quad R_{CM} = \frac{M_{Moon} r + M_E \times 0}{M_E + M_{Moon}}$$

$$= 4.7 \times 10^3 \text{ km}$$

$$\Rightarrow M_{Moon} = M_E \frac{R_{CM}}{r - R_{CM}}$$



2) Consider two planets with masses M_1 and M_2 , and radii R_1 and R_2 , sitting at $x_1 = 0$ and $x_2 = r$. For the first two parts assume they are not moving.

a) Find the point between these two masses that the gravitational field is zero and call it X .

b) We throw a stone from planet 2 to planet 1 from point P . What is the minimum initial velocity, v_i , needed to do that? We only need to reach the point you found in the previous part.

Now we assume that $M_1 \gg M_2$, and M_2 is rotating around M_1 in a circular orbit. You can think about sun-earth or earth-moon for example, but there is no spin. It turned out that to throw a stone with minimum push, from planet two to planet one, you should throw it as shown in fig. 1, using Hohmann path. We throw the stone, with velocity v relative to M_2 , in the direction that M_2 moves so it adds to the velocity. This path is an ellipse which you know the major axis. It touches the surface of planet 1 and stops.

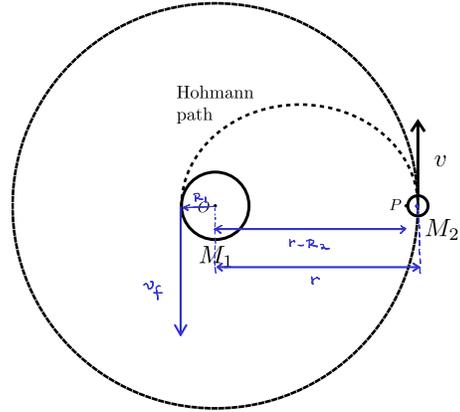


Figure 1: Throwing stone using Hohmann path.

c) What is the velocity of M_2 around M_1 ? Call this V , and define $v_i = v + V$.

d) Angular momentum of the stone with respect to point O is conserved in this path. Write down an equation for this in terms of v_i, v_f, r, R_1, R_2 .

e) How long it takes for the stone to travel this path? ($r \gg R_2$)

f) Using the fact that the energy of the stone is conserved through the path and previous parts, find v_i .

a) $g_1 = g_2 \Rightarrow \frac{GM_1}{x^2} = \frac{GM_2}{(r-x)^2}$

$x \neq 0, r \Rightarrow M_1(r-x) = M_2 x^2$

$\Rightarrow \sqrt{M_1} |r-x| = \sqrt{M_2} |x| \quad 0 < x < r \Rightarrow x = r \frac{\sqrt{M_1}}{\sqrt{M_1} + \sqrt{M_2}} = r \left[1 + \sqrt{\frac{M_2}{M_1}} \right]^{-1}$

M_1 at $x=0$, M_2 at $x=r$, $x=X$

b) at point $x=X$, $v=0$, & the energy is conserved, (m is mass of the stone)

$E = -\frac{GM_1 m}{x} - \frac{GM_2 m}{r-x}$, right after throwing at P , $v = v_i$ and the stone is at the surface of M_2 , so $E = \frac{1}{2} m v_i^2 - \frac{GM_1 m}{r-R_2} - \frac{GM_2 m}{R_2}$. So $v_i^2 = 2G \left[\frac{M_1}{r-R_2} - \frac{M_1}{x} + \frac{M_2}{R_2} - \frac{M_2}{r-x} \right]$.

if $r, x \gg R_1, R_2$ then $v_i^2 = 2G \frac{M_2}{R_2}$, which is the escape velocity from M_2 if M_1 was not there.

c) $F = ma : \frac{GM_1 M_2}{r^2} = M_2 \frac{V^2}{r} \Rightarrow V = \sqrt{\frac{GM_1}{r}}$

d) at the beginning, $L_i = m r v_i = m R_2 (v + V)$, at the final point, $L_f = m R_1 v_f$, so $r v_i = R_1 v_f$.

e) for the ellipse, the major axis is, $2a = r + R_1$, and we need almost half the period ($r \gg R_2$), so $\frac{T}{2} = \frac{1}{2} \left[\frac{4\pi^3 a^3}{GM_1} \right]^{1/2} = \frac{\pi (r+R_1)^{3/2}}{2\sqrt{GM_1}}$.

f) $E = \frac{1}{2} m v_i^2 - \frac{GM_2 m}{R_2} - \frac{GM_1 m}{r} = \frac{1}{2} m v_f^2 - \frac{GM_1 m}{R_1} - \frac{GM_2 m}{r+R_1}$ & $r v_i = R_1 v_f$, so $v_i = \left[\frac{2G \left(\frac{M_2}{R_2} - \frac{M_1}{r} + \frac{M_1}{R_1} - \frac{M_2}{r+R_1} \right)}{\left(\frac{r}{R_1} \right)^2 - 1} \right]^{1/2}$

and $v = v_i - V = v_i - \sqrt{\frac{GM_1}{r}}$, the velocity we need at the beginning.